## svllabC25-2022.txt Jan 11, 22 23:46 Page 1/2 SYLLABUS for ASTC25, Winter 2022. Title: Astrophysics of Planetary Systems Lecturer: Prof. Pawel Artymowicz (pronounced PAvel ArtyMOvich) URL for all course materials: http://planets.utsc.utoronto.ca/~pawel/ASTC25 Tutorials: Thursdays 12-13 [in HW 408 from Feb?] online zoom (no tutorial on da y 1) Lectures: Thursdays 13-15 [in HW 408 from Feb?] online zoom meetings Assignments are due at 12:00 (beginning of tutorial). Submit via Quercus. Calendar of Lectures (L), Tutorials (T), Assignments (A): 13 Jan -- L1+2 20 Jan T1 L3+4 27 Jan T2 L5+6 3 Feb T3 L7+8 10 Feb T4 L8 17 Feb T5 L9 24 Feb -- (reading week) 3 Mar T6 L10-11 (T6 = midterm, time: 12-13) 10 Mar T7 T.12-14 17 Mar T8 L15+16 A3 24 Mar -- L17-20 31 Mar T9 L21+22 7 Apr T10 L23+24 A4 April, FINAL EXAM: TBD Syllabus is subject to ongoing small changes. Please download weekly. Numbers in brackets = chapt. of the Lissauer-dePater textbook to read, preferably ahead of the lecture. 1. Introduction to the subject and style of the course 2-4. Gravitational mechanics of planetary systems \* Gravitational 2-body interaction \* Kepler's laws with derivations \* The 2-body problem and the elliptic motion: E. L. vs. a. e 5. Elements of celestial mechanics I \* Tides in the solar system \* Disruption of satellites: the Roche limit 6-7. Elements of celestial mechanics II \* Precession of orbits and spin axes \* Theory of perturbations vs. numerical computations \* Restricted 3-body problem and the Hill problem 8. Orbits beyond the elliptic ones \* Stability of motion \* Lagrange points and in disk \* Orbital resonances and chaos \* The future of our solar system 9-10. Formation of disks and stars (ch. 15) \* Giant molecular clouds \* Jeans instability of protostellar cloud cores \* Opacity-limited fragmentation \* Simulations: the ubiquity of protostellar disks, brown dwarfs 11-12. Origins: Accretion disks (ch. 15) \* Analogue disks: AGN/quasar disks, and their accretion \* Accretion disk geometry \* Disks as evolving, shearing flows 13. Formation od planets: the main scenarios \* Accumulation versus fragmentation: scenarios for the giants \* Gravitational stability of protoplanetary disks \* From dust to planetesimals 14. Formation od planets: early stages

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* Gravitational scatteri		ing
* Late heavy bombardment * Core-instability and o	gas accretion onto giant planets	
<ul><li>* Interplanetary dust: 2</li><li>* Vega-type systems, rep</li></ul>	solar and extrasolar systems Zodiacal light disk and Brownlee part plenished dusty disks of planetary sy vidence of planetesimals and planets t avalanches	
16. Asteroids, planetoids (c    * Kirkwood gaps in aster    * Asteroid belt    * Kuiper belt objects ar    * Planetoids/dwarf plane    * Classification of meter	roid belt nd mechanics ets: Eris and others	
17. Water in planetary syste  * The H-comets  * Comets - icy dirtballs  * Reservoirs: Oort clouc  * Halley, Hyakutake, Hal  * Where do Earth's ocean	s or dirty iceballs? d, inner Oort cloud, Kuiper belt le-Bopp, Holmes2	
* Examples of satellites	s launching waves at resonances	
18. Irradiation instability  * Dust avalanches  * Role of optical thickr  * Numerical results	-	
19-20. Exoplanet discovery  * Methods: timing, radia  * Overview of results ar  * chemical correlations	al vel., transits, microlensing, imag nd examples	ing
<pre>21. Atmospheres of planets a   * Equilibrium of atmosph   * Greenhouse effect   * Effect of changing alk</pre>	heres	
22-23. Migration of protopla  * Migration as the reas  * Type 0 - by gas drag  * Type I - Lindblad torc  * Type II - after gap op  * Type III - partially of	on for diversity ques pening	
24. Astrobiology and SETI  * Life on Earth: local of  * Life elswhere: Mars, F  * Habitable zones  * Drake's equation, SETI	Europa, moons of exoplanets?	